



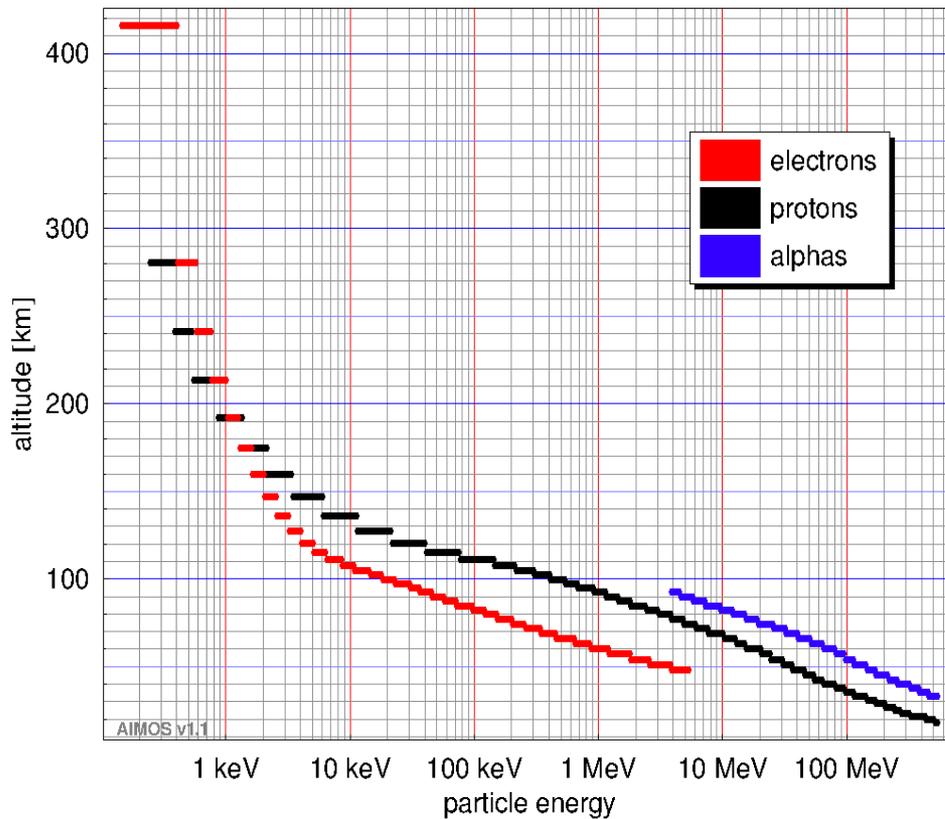
## **Modeling ionosphere response to Solar Proton Events in the whole atmosphere model EAGLE**

F.S. Bessarab<sup>1</sup>, T.V. Sukhodolov<sup>1,2,3</sup>, M.V. Klimenko<sup>1</sup>,  
V.V. Klimenko<sup>1</sup>, Yu.N. Korenkov<sup>1</sup>, B. Funke<sup>4</sup>, I.E.  
Zakharenkova, E.V. Rozanov<sup>1,2,3</sup>

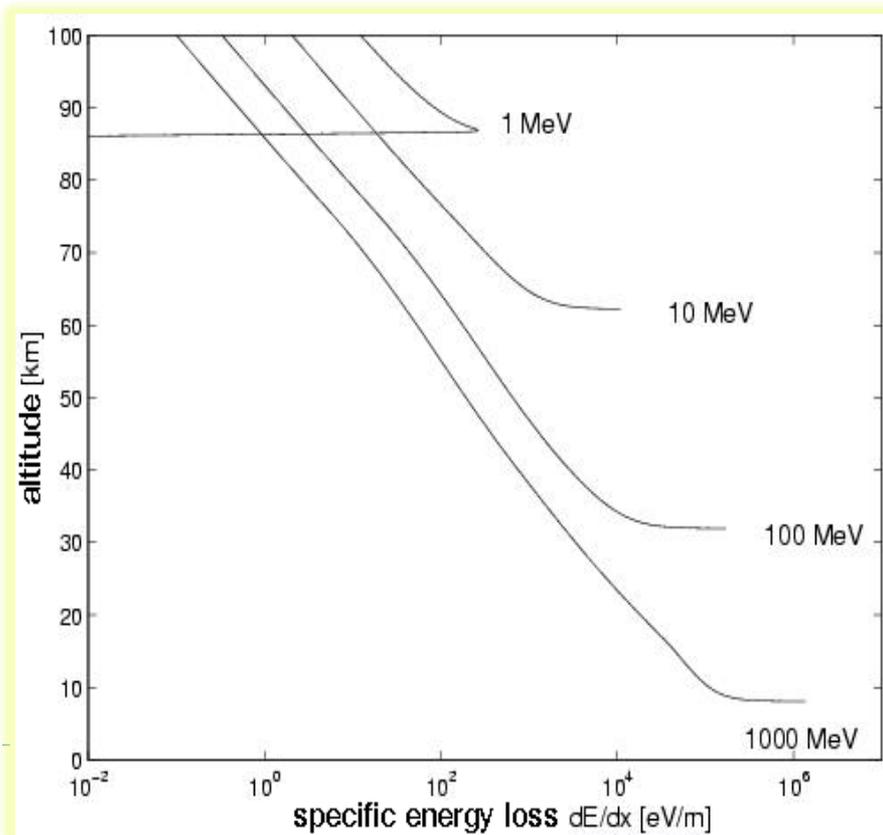
1. WD IZMIRAN, Kaliningrad, Russia
2. PMOD/WRC, Davos, Switzerland
3. IAC ETH Zurich, Switzerland
4. IAA-CSIC, Granada, Spain

# Altitude of maximum energy deposition

(Wissing and Kallenrode 2009)



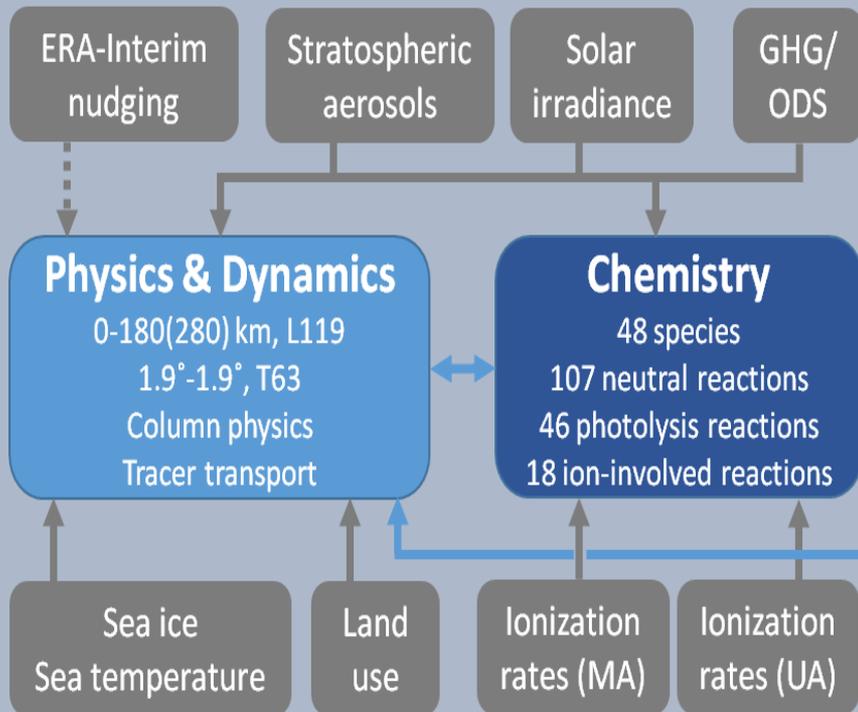
(Quack, 2005)



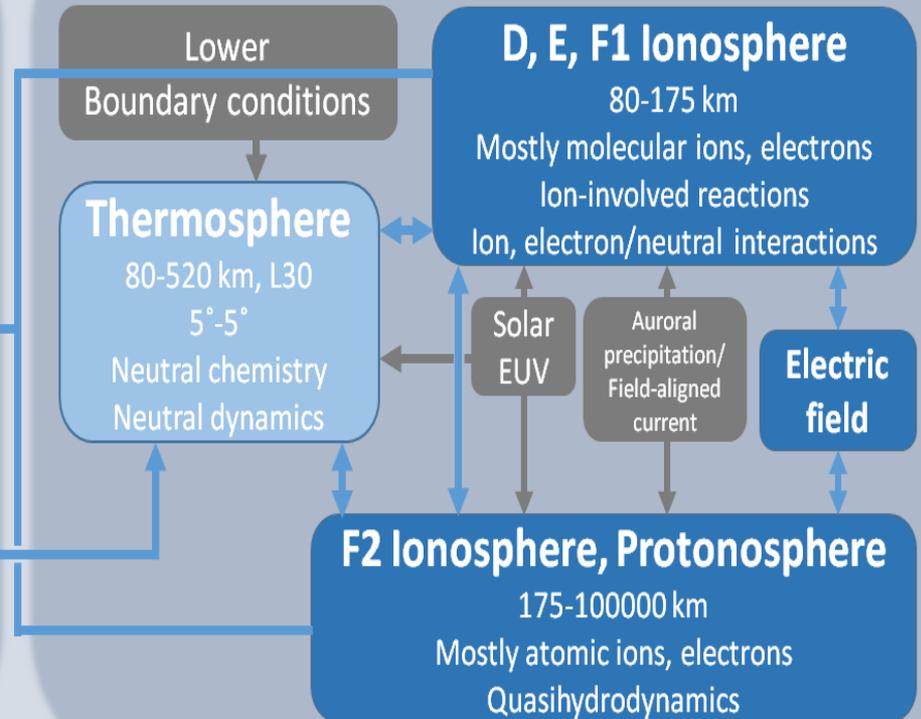
# Whole Atmosphere Global Model

## EAGLE

### HAMMONIA



### GSMTIP



# Atmospheric Ionization Module Osnabrück (AIMOS)

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AIMOS models of ion pair production due to precipitating particles solar and

AIMOS is designed to and magnetospheric 3-D ionization pattern

AIMOS consists of two simulation and a set of two polar-orbiting satellites depending on geomagnetic activity

**The POES satellites** (POES: Polar Orbiting Environmental Satellite) are polar-orbiting satellites in a Sun synchronous orbit with a height of 850 km and an inclination of 98°

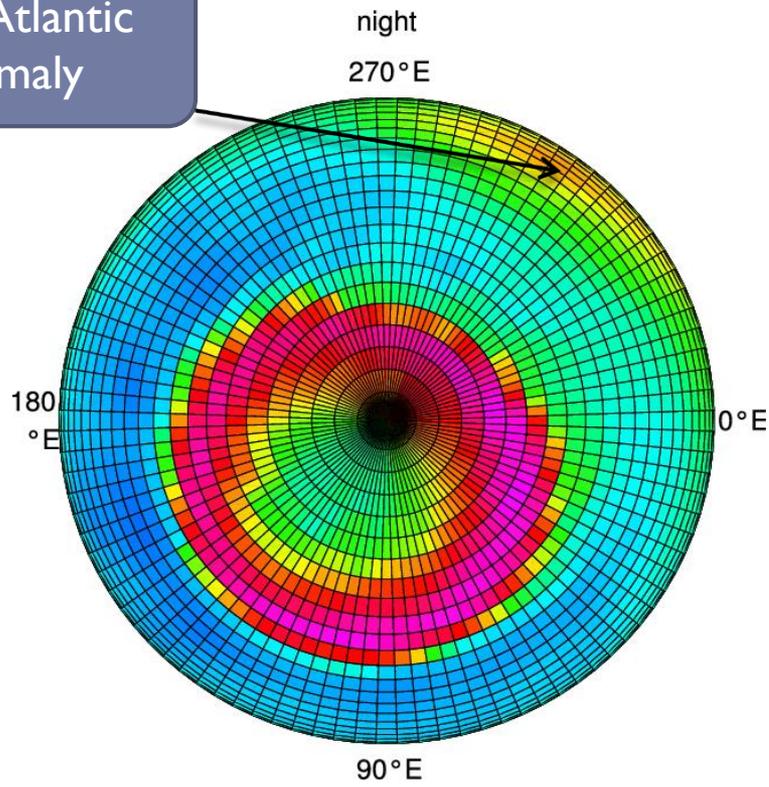
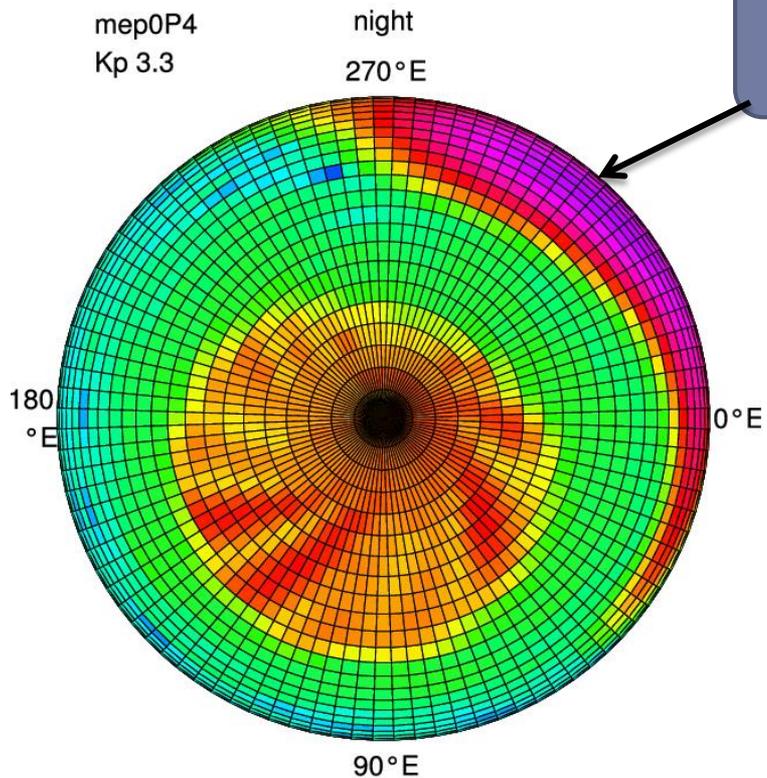
**The GOES satellites** (GOES: Geostationary Operational Environmental Satellite) are in geostationary orbit located at W135 and W104.

# AIMOS

Higher energetic particles  
(0.8– 2.5 MeV protons) flux.

Low energetic particles (30–80  
keV protons) flux

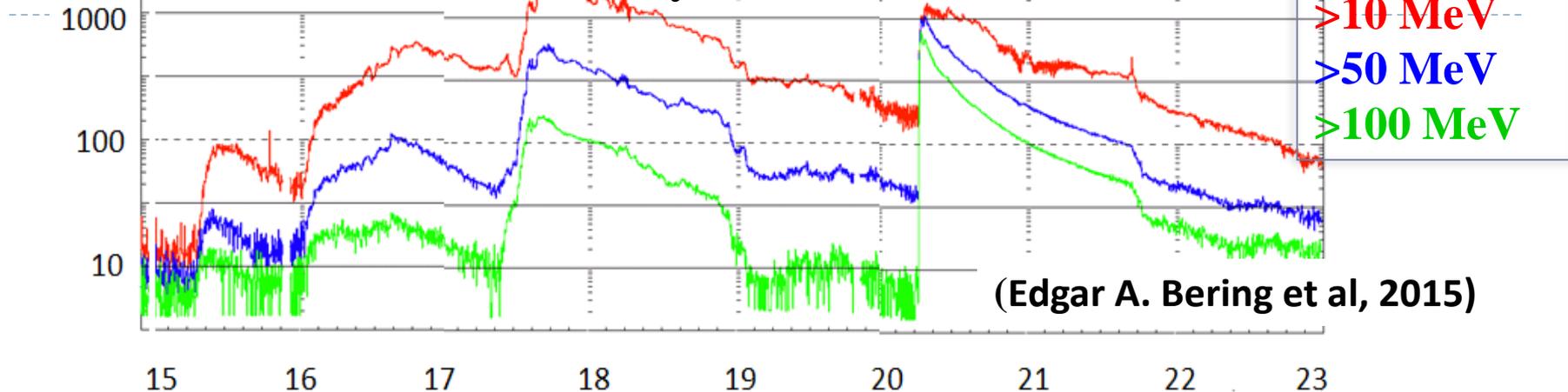
South Atlantic  
Anomaly



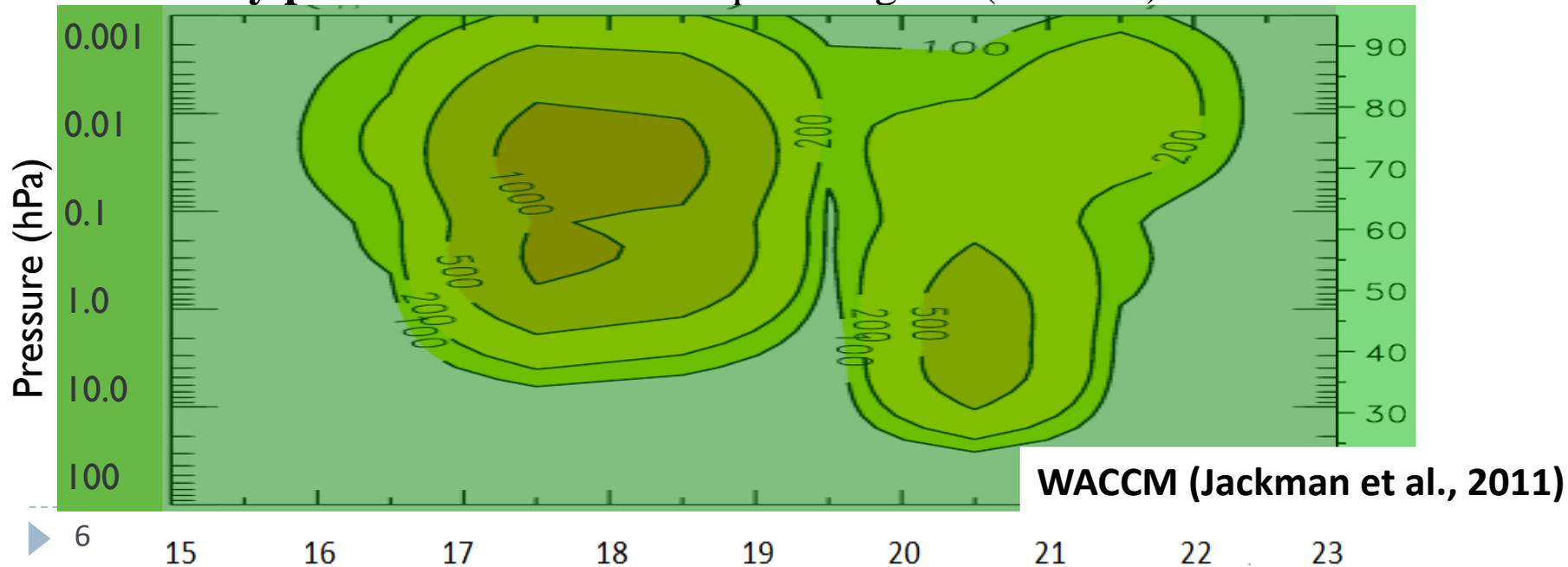
A view on the south pole  
is shown.

# Solar Proton Events, January of 2005

Proton Flux Measurements by GOES 11 ( $\text{cm}^{-2} \text{s}^{-1} \text{sr}^{-1}$ )



Ion rates by protons in the northern polar region ( $\text{cm}^{-3} \text{s}^{-1}$ )



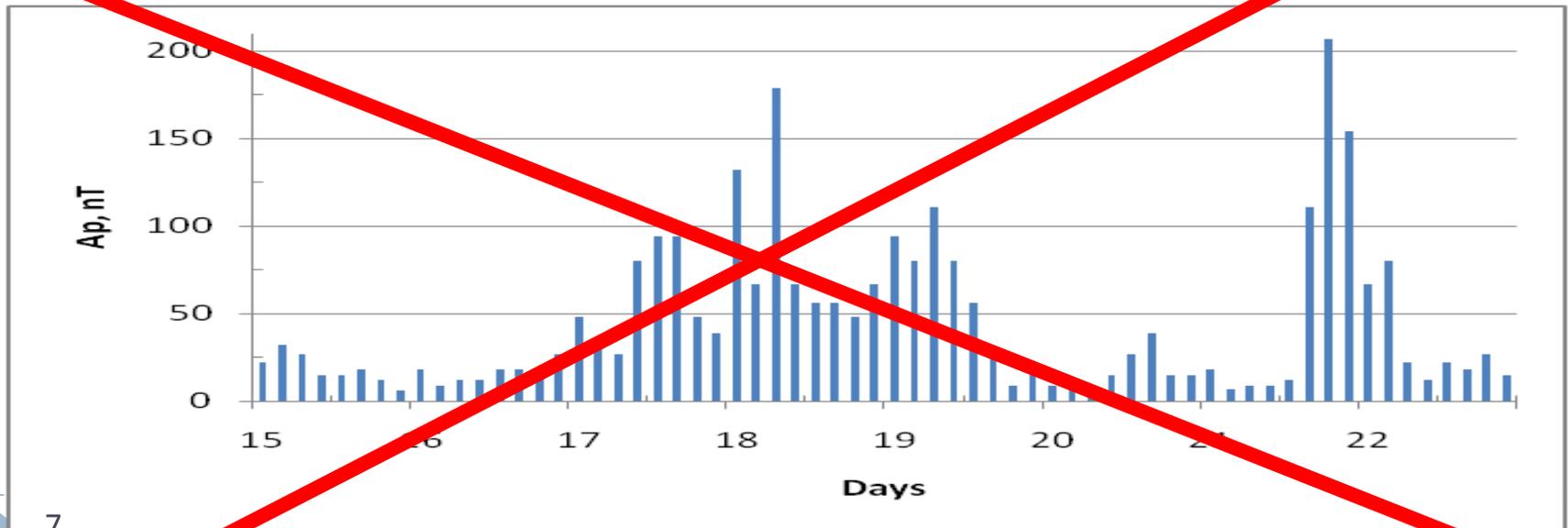
# EAGLE simulations

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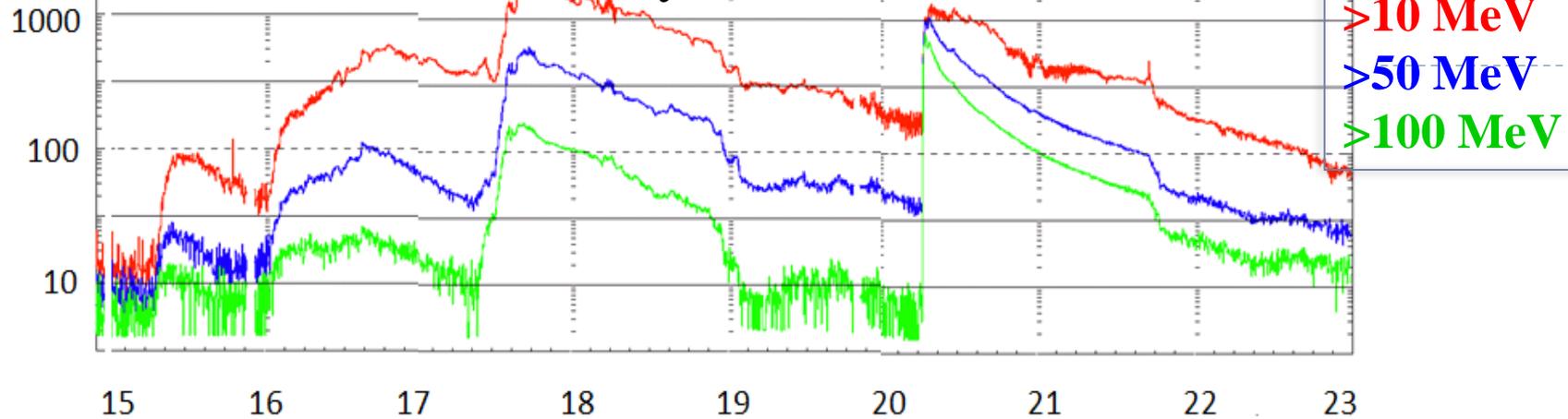
Data comparisons for two model runs:

- reference model simulation (quiet condition)
- SPE simulation.

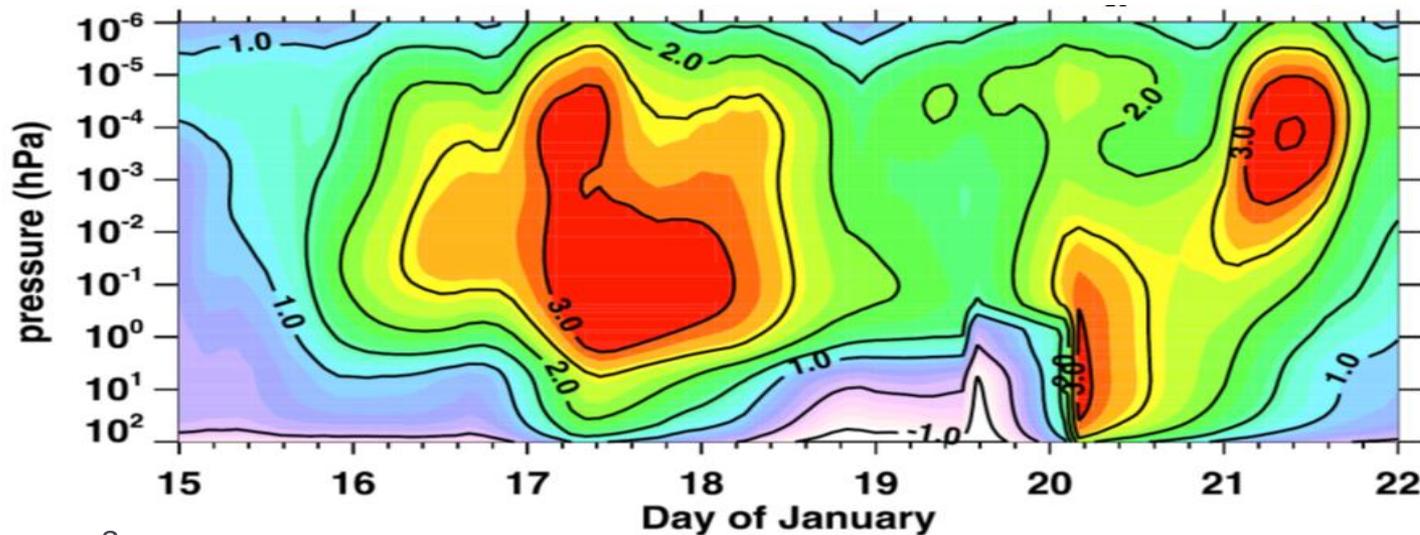
It was changed **ONLY** proton flux in AIMOS



# Proton Flux Measurements by GOES 11 ( $\text{cm}^{-2} \text{s}^{-1} \text{sr}^{-1}$ )

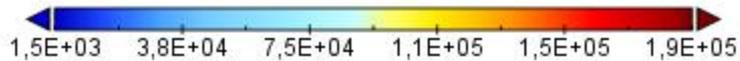
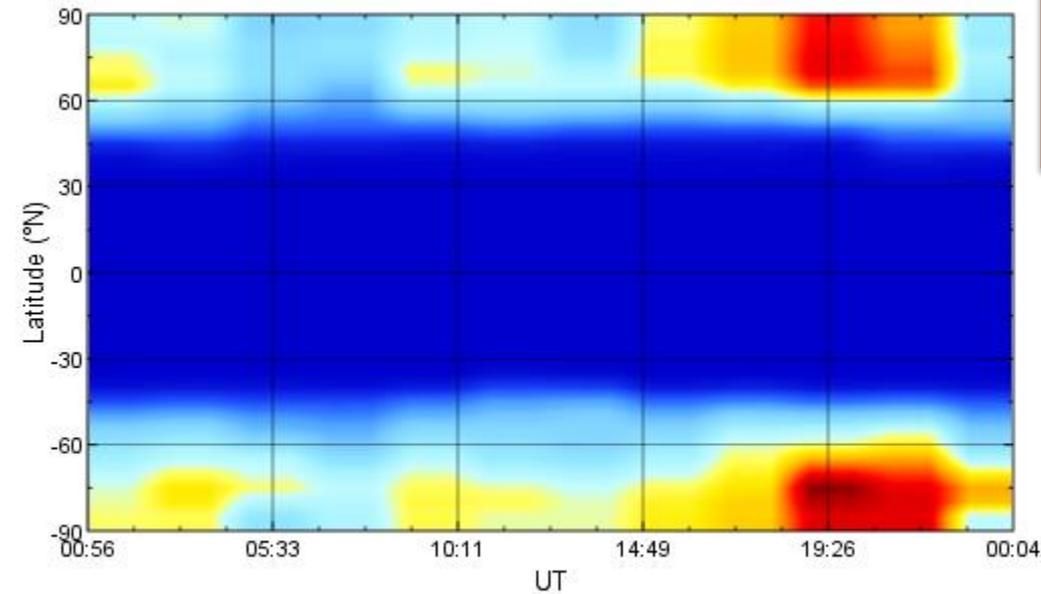


## EAGLE log-ionization rates by protons (AIMOS) on 80 N ( $\text{ion pairs cm}^{-3} \text{s}^{-1}$ )

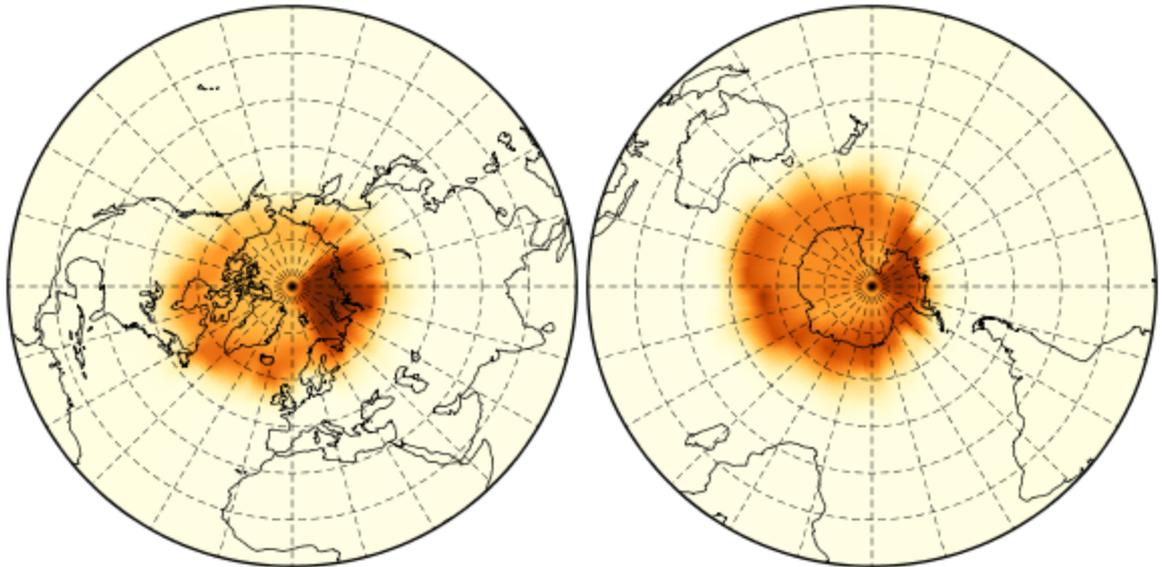


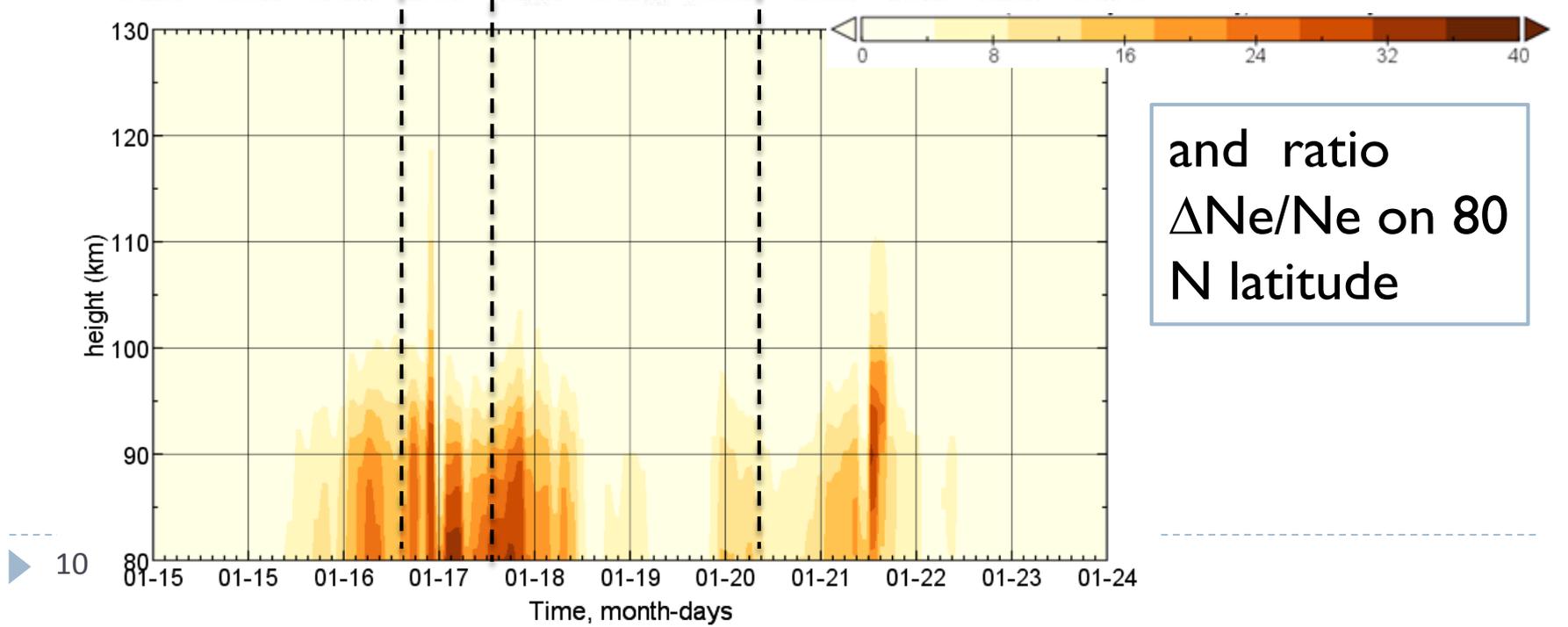
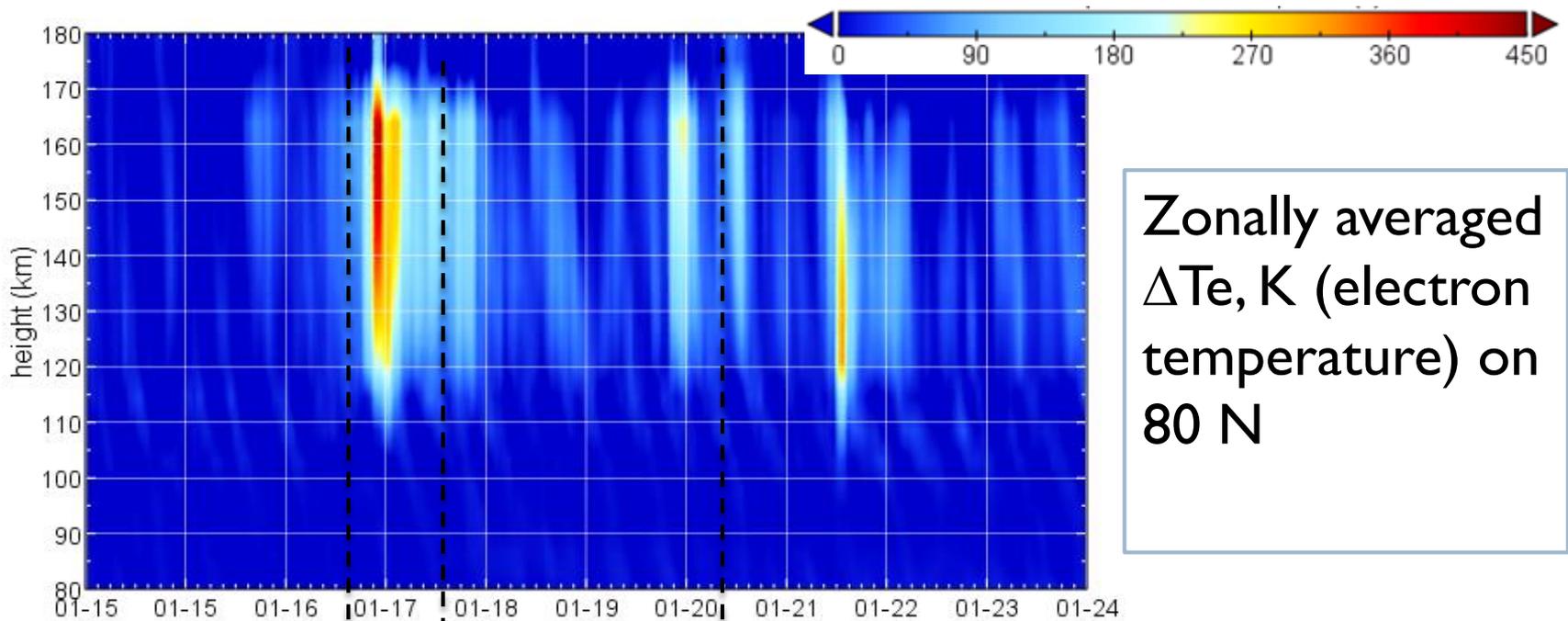
# SOLAR PROTON EVENT, EAGLE model simulation

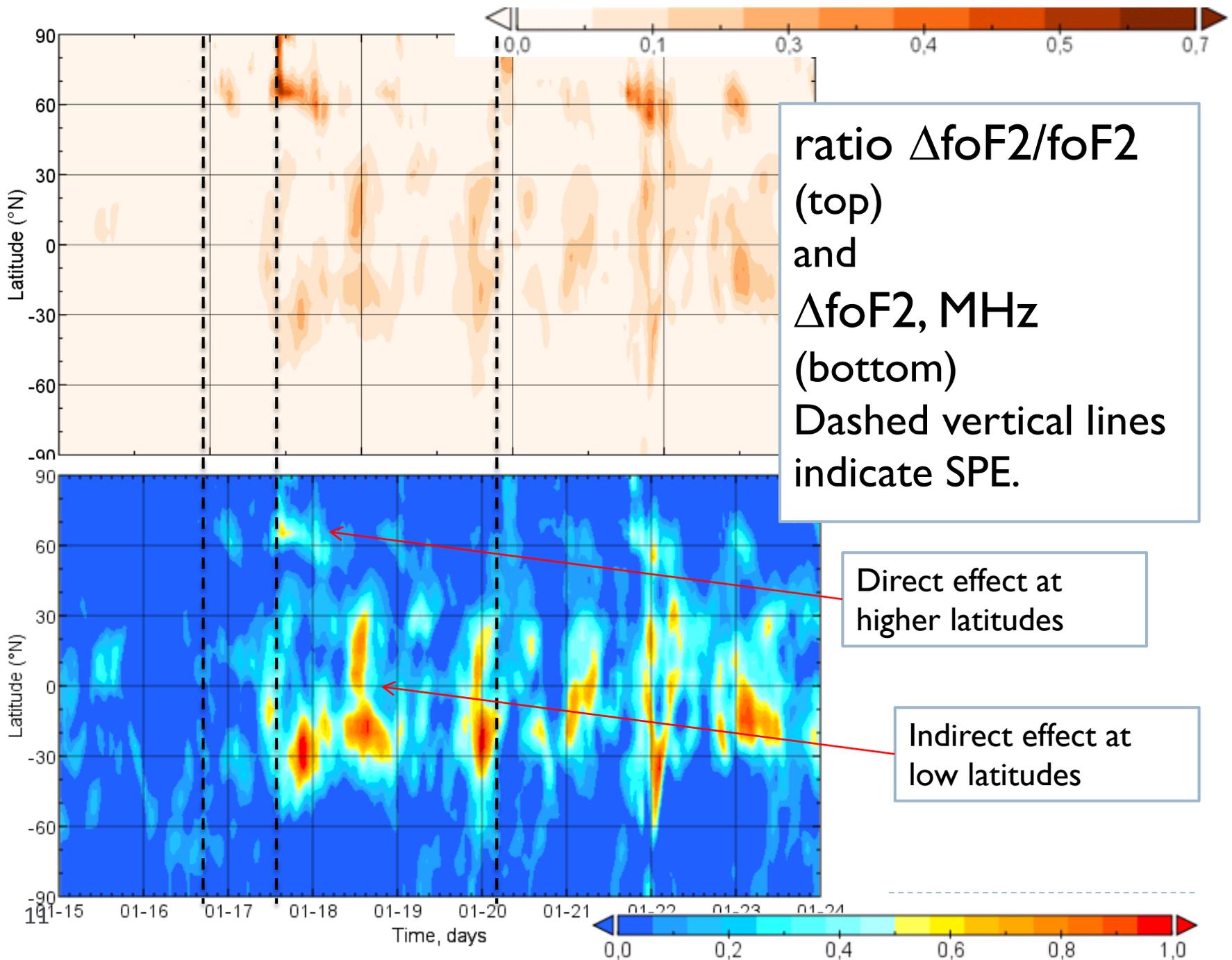
Zonally averaged Ne,cm-3  
at 83 km, January, 17

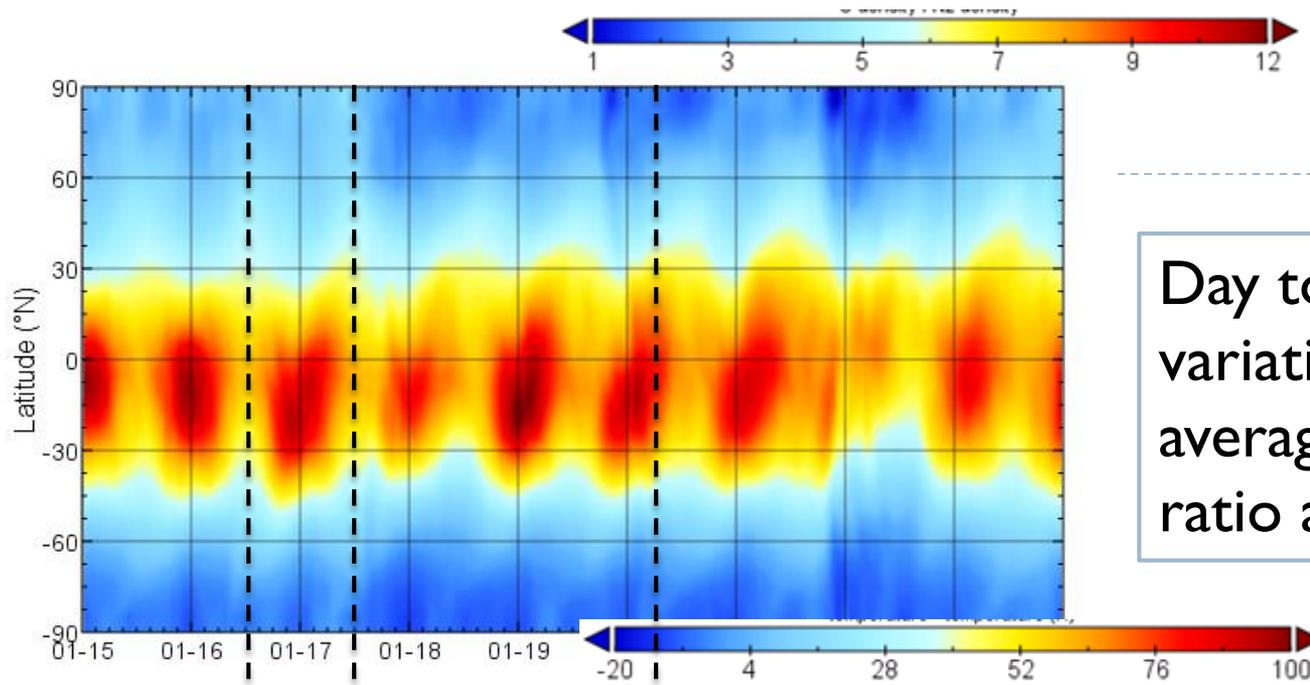


Ne,cm-3 at 83 km,  
17 of January,  
19 UT

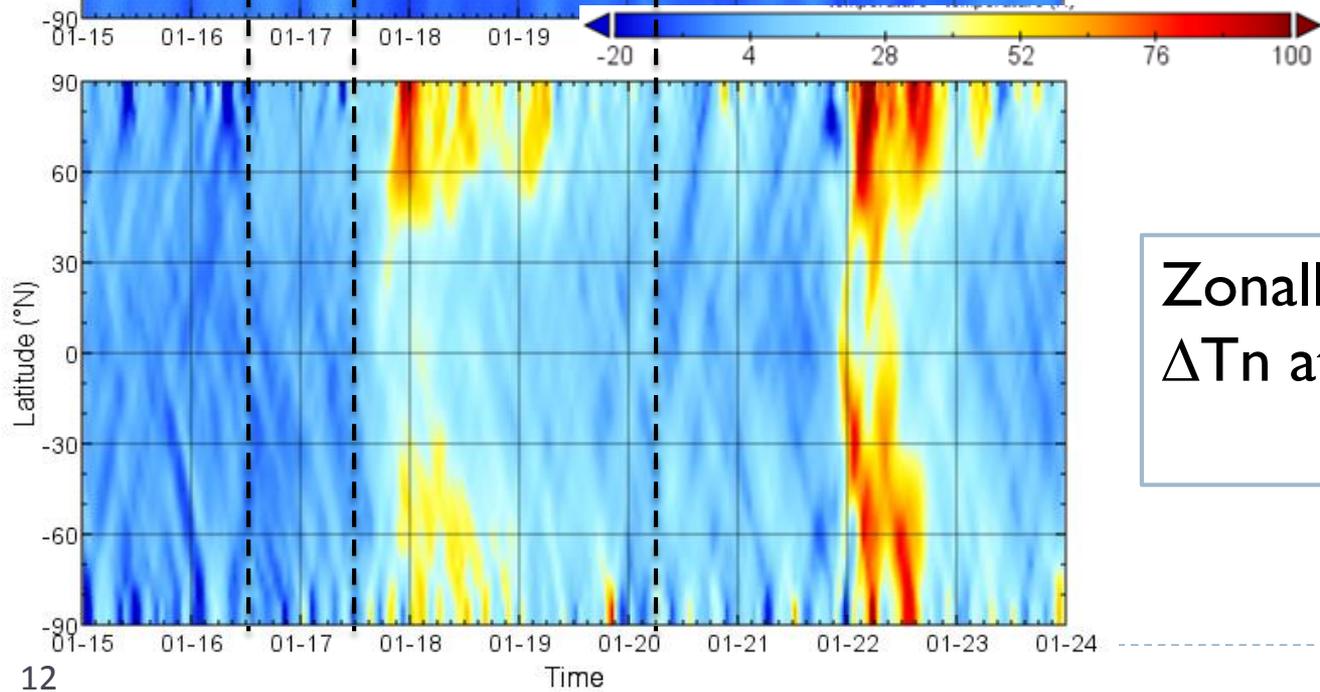






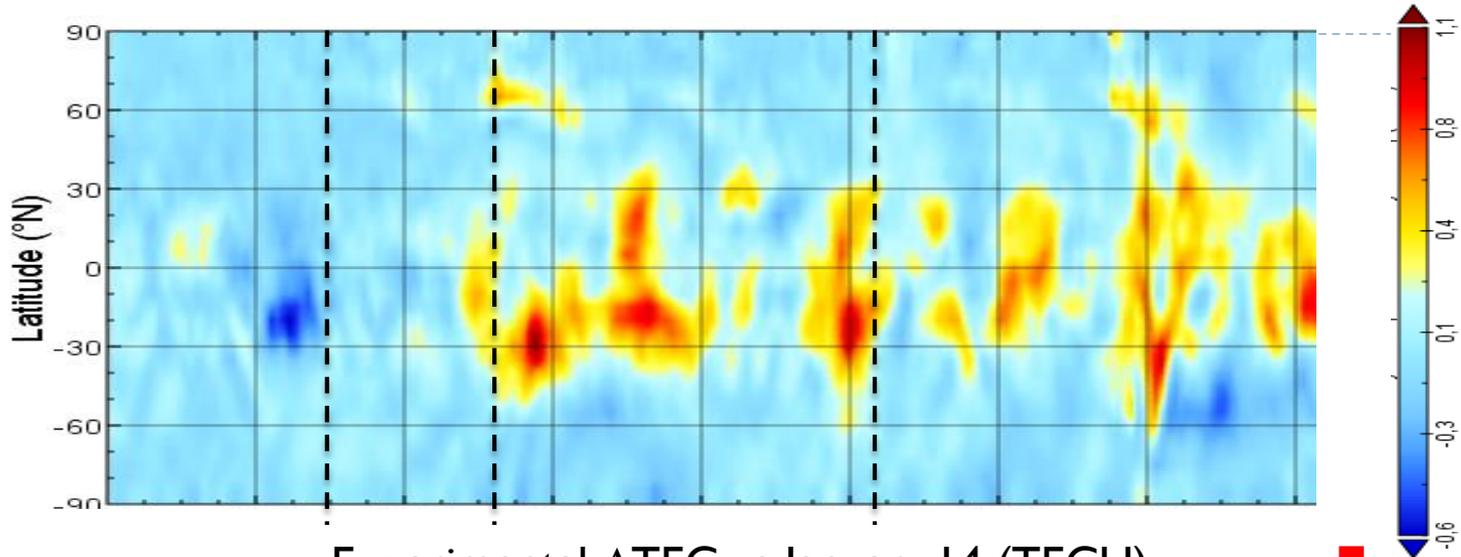


Day to day  
variation zonally  
averaged O/N<sub>2</sub>  
ratio at 300 km

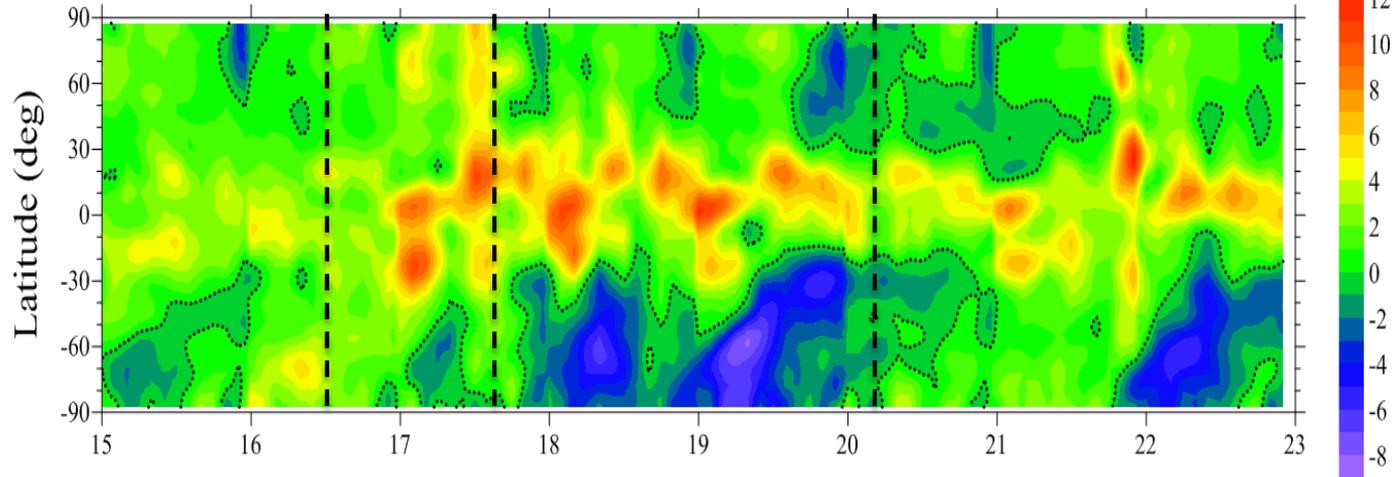


Zonally averaged  
 $\Delta T_n$  at 300 km

# EAGLE $\Delta\text{foF2}$ ( $\text{foF2}_{\text{spe}} - \text{foF2}_{\text{ref}}$ ), MHz



# Experimental $\Delta\text{TEC}$ vs January, 14 (TECU)



January 2005

# Summary

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- The whole atmosphere model (EAGLE) allows to investigate large-scale phenomena covering all layers of the atmosphere.
- The response of the ionosphere to the proton event of January 2005 is considered.
- Despite the relative transparency of the thermosphere to the high-energy particles, ionospheric response on SPE was obtained.
- The direct and indirect effects of the proton event due to ionization and variations of the atmosphere parameters, respectively, should be noted.

**The study was supported by RSF grant No. 17-17-01060**